Resonant electron-molecular cation collisions in the edge plasmas of fusion devices: new state-to-state cross sections and rate coefficients







Joint ICTP-IAEA School and Workshop on Fundamental Methods for Atomic, Molecular and Materials Properties in Plasma Environments, April 16-20 2018, Trieste Italy.





Dissociative recombination





Vibrational excitation





Dissociative excitation









Introduction: Cold ionized media

Interstellar molecular clouds

Planetary atmospheres

Cold laboratory plasmas

At the wall of the fusion devices (ITER) project









Hypersonic entry of spacecrafts



Plasma-assisted depollution





Broadband Emission (no bias)



Plasma-assistedcombustion



Theoretical approach: Reactive collisions

$$e^{-} + AB^{+}(v_{i}^{+}, N_{i}^{+}) \to AB^{*}, AB^{**} \to \begin{cases} AB^{+}(v_{f}^{+}, N_{f}^{+}) + e^{-} & \text{EC,IC,SEC} \\ [A + B]_{(\epsilon)} & DR \\ A^{+} + B + e^{-} & DE \end{cases}$$

Poster of J. Zs. Mezei *et al*

Theoretical approach: MQDT

Multichannel Quantum Defect Theory

Seaton (1958-1983), Fano, Jungen, Greene, Giusti -Suzor (1970-...),...



MQDT: DR mechanisms



Case study: SH⁺

K. M. Menten et al: Submillimeter absorption from SH+, a new widespread interstellar radical, 13CH+ and HCI, A&A 525, A77 (2011)

c-C₃H₂ $H^{13}CO^{+}/5.9~10^{11}$ d 2 c-C₃H₂ [[h k/3] THE JOURNAL OF CHEMICAL PHYSICS 146, 204109 (2017)



Discovery: Benz et al (2010)



Formation and destruction routes $S^++H_2 \rightarrow SH^++H (-0.86 \text{ eV})$

A theoretical study of the dissociative recombination of SH⁺ with electrons through the ²Π states of SH

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(Received 15 February 2017; accepted 5 May 2017; published online 30 May 2017)

A quantitative theoretical study of the dissociative recombination of SH⁺ with electrons has been carried out. Multireference, configuration interaction calculations were used to determine accurate potential energy curves for SH⁺ and SH. The block diagonalization method was used to disentangle strongly interacting SH valence and Rydberg states and to construct a diabatic Hamiltonian whose diagonal matrix elements provide the diabatic potential energy curves. The off-diagonal elements are related to the electronic valence-Rydberg couplings. Cross sections and rate coefficients for the dissociative recombination reaction were calculated with a stepwise version of the multichannel quantum defect theory, using the molecular data provided by the block diagonalization method. The calculated rates are compared with the most recent measurements performed on the ion Test Storage Ring (TSR) in Heidelberg, Germany. Published by AIP Publishing. [http://dx.doi.org/10.1063/1.4983690]

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Received 28 January 2014 / Accepted 19 July 2014

Abstract

Context. Tens of light hydrides and small molecules have now been detected over several hundreds sightlines sampling the diffuse interstellar medium (ISM) in both the Solar neighbourhood and the inner Galactic disk.

Aims. These new data confirm the limitations of the traditional chemical pathways driven by the UV photons and the cosmic rays (CR) and the need for additional energy sources, such as turbulent dissipation, to open highly endoenergetic formation routes. The goal of the present paper is to further investigate the link between specific species and the properties of the turbulent cascade in particular its space-time intermittency.

Methods. We have analysed ten different atomic and molecular species in the framework of the updated model of turbulent dissipation regions (TDR). We study the influence on the abundances of these species of parameters specific to chemistry (density, UV field, and CR ionisation rate) and those linked to turbulence (the average turbulent dissipation rate, the dissipation timescale, and the ion-neutral velocity drift in the regions of dissipation).

Results. The most sensitive tracers of turbulent dissipation are the abundances of CH⁺ and SH⁺, and the column densities of the J = 3, 4, 5 rotational levels of H₂. The abundances of CO, HCO⁺, and the intensity of the 158 μ m [CII] emission line are significantly

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MQDT: molecular data



GAMESS + Block diagonalization method (Kashinki, Hickman, Talbi)

































SH⁺: DR rate
$$\alpha(T) = \left(\frac{m_e}{2\pi kT}\right)^{3/2} \int_0^\infty \sigma(v)v \exp\left(-\frac{m_e v^2}{2kT}\right) 4\pi v^2 dv$$



Results: The most abundent molecule





$$H_2$$



Densities ~ 10^{5} - 10^{-7} cm⁻³ T ~ 30000 K - 0.003 K

Warm ionized medium (densities $\sim 0.3 \text{ cm}^{-3} - T \sim 10000 - 8000 \text{ K})$ Warm neutral medium (densities $\sim 0.3 \text{ cm}^{-3} - T \sim 8000 \text{ K})$ Cold neutral medium (densities $\sim 30 \text{ cm}^{-3} - T \sim 50 \text{ K})$ Molecular clouds (densities $> 100 \text{ cm}^{-3} - T > 10 \text{ K})$ Some orders of magnitude : * $n_e = 10^8 \cdot 10^{12} \text{ cm}^{-3} (<10^{-2} \text{ and more often } <10^{-5})$ * $<\varepsilon_e> = 1-10 \text{ eV}$ * $T_g = 300 - 6000 \text{ K}$ * $'T_v' = 1000 - 5000 \text{ K} \text{ (molecular gases)}$

Results: The most abundant molecule *a low temperatures*

Rotational excitation



Dissociative recombination



Reactive collisions of very low-energy electrons with H_2^+ : rotational transitions and dissociative recombination

M. D. Epée Epée,¹ J. Zs Mezei,^{2,3,4} O. Motapon,^{1,5}* N. Pop⁶ and I. F. Schneider^{2,3}*

and its isotopoloques

Results: The most abundant molecule *a medium temperatures*



in progress for its isotopoloques

Results: The most abundant molecule (a) high temperatures



Dissociative recombination of electrons with diatomic molecular cations above dissociation threshold: Application to H₂⁺ and HD⁺

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in progress for its isotopoloques

Results: Molecules in fusion experiments





@ low temperatures

State-to-state chemistry and rotational excitation of CH⁺ in photon-dominated regions

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Conclusions

MQDT: state-to-state calculations

 Temporary captures into superexcited states: HUGE RESONANT **EFFECTS**

Data needs: di- and poly-atomics



 $- = C_1 \& C_2 - D_1 \& E_1 \\ \dots C_1 - D_1$

5pg.v=3R(0)

9pπ,v=2 Q(1)

9pg.v=2 R(1)



E (eV)

In collaboration with

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- N. Pop (Politech. Univ. Timisoara), F. Iacob (West Univ. Timisoara), O. Motapon, M. D. Epee Epee (Univ. Douala), S. Niyonzima (Univ. of Burundi)
- J. Tennyson, D. A. Little (Univ. College London), K. Chakrabarti (Univ. of Kolkatta)
- D. Talbi (Univ. Montpellier), D. O. Kashinski (West Point), A. P. Hickman (Lehigh Univ.)
- Ch. Jungen, J. Robert, O. Dulieu (Univ. Paris Sud)
- Å. Larson (Stockholm Univ.), A. E. Orel (Univ. of California Davis), V. Kokoouline (Univ. of Central Florida)





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SUperexcited MOlecular STates of Astrophysical Importance : dissociative recombination and spectroscopy





Thank you for the attention!





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