## Reliable atomic data for fusion research and 'astrophysics:

Benchmarking calculations for highly charged ions



für Kernphysik

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## HCI as resource for fundamental physics

- Atoms are key assets in fundamental studies:
  - universally reproducible
  - spectroscopic frequency references from MW to VUV
  - mass references for nuclear, neutrino physics, etc.
  - EDM searches for interactions beyond the SM
  - atomic interferometric probes of space-time, relativity
- Limitation: Unstable at high interaction energies
- Using HCI expands these possibilities:
  - high intrinsic stability against perturbations
  - stable at high photon energies up to the 0.1 MeV
  - in greater variety than atoms or singly charged ions

Review of Modern Physics coming up: (arXiv, Kozlov, Safronova, JRC, Schmidt 2018)



In the Universe, elements are mostly highly ionized: Highly charged ions (HCI)

- Interior of the Sun (15 MK)
- Solar corona (2 MK)
- Solar wind (MK)
- Supernova remnants
- Active galactic nuclei (100 MK)
- Warm-hot intergalactic medium (0.1-1 MK)

#### In the laboratory:

- Fusion machines (50 MK)
- Accelerators, laser produced plasmas (1 MK)
- Electron beam ion traps (e.g. in Heidelberg)









## State of the art in the field of HCI

- X-ray photon energies
- VUV photon energies
- Optical photon energies
- Lifetimes (ns... ms)
- Natural linewidths X-rays:

Homo Heiðelbergensis

1.5 ppm 4 ppm 0.3 ppm 0.15 % resolved

Accuracy is 12 orders of magnitude lower than in frequency metrology

Stone-age spectroscopy at the 10<sup>-6</sup> level



#### HCI: Under way in the "spectral desert"

- No reports about the ions of interest and no transition data available for most ions!
- HCI production in EBIT easy, identification much harder





## EBIT facility at MPIK





## Overview

- Electron beam ion traps
- Photon excitation studies
- Electron excitation studies
- Optical and EUV spectroscopy
- Charge exchange
- Frequency metrology

#### HCI production with electron beam ion trap



Electron beam drives ionization, excites and traps the ions inside a cylindrical volume

#### Electron beam ion traps big and small



- Unique facility at MPIK, supporting Pfeifer and Blaum division
- Out of ~20 research EBITs worldwide, 10 are at or come from MPIK

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## EBIT facility at MPIK





### Some compac EBITs









## Electron beam ion traps at MPIK

#### Cryogenic devices:

HD EBIT: stationary machine built in 1999 (former FreEBIT)
FLASH EBIT: transportable for beamtimes at FLASH, BESSY, LCLS and PETRAIII
Hyper EBIT

Miniature:

Polar-X EBIT (at PETRAIII)
Tip-EBIT (Blaum division)
TT-EBIT (Blaum division)
CryPTEx-II-EBIT









# Strong excitation by electron impact and resonant photorecombination



## Ge, Si detectors, crystal and grating spectrometers, microcalorimeters, etc., for X-ray diagnostics



# Choice of charge state by dialing electron beam energy





## What we do with EBITs

- Photorecombination processes: radiative + dielectronic, trielectronic and quadruelectronic recombination,
- Photoionization of HCI with synchrotron radiation, from N<sup>3+</sup> (from 60 eV) to Kr<sup>33+</sup> (at 14 keV)
- High-resolution spectroscopy from optical to X-rays
- Free-electron laser soft x-ray spectroscopy (<800 eV)
- High-resolution x-ray metrology with synchrotron radiation (<14.4 keV)
- Laser spectroscopy of forbidden optical lines in HCI
- Sympathetic cooling of HCI for frequency metrology
- Charge-exchange studies



## Free-electron laser and synchrotron-radiation excitation and photoionization of highly charged ions



## Photonic interactions with HCI

- Combine novel X-ray sources (free-electron lasers, synchrotrons) with electron beam ion trap (EBIT)
- to measure excitation energies beyond current accuracy limits
- to determine cross sections and line profiles for photoexcitation, photoionization HCI benchmarking atomic theory



## Interaction with photon beams



cryostat
 superconducting magnet
 electron gun

trapped ions

#### photon beam

#### drift tubes

#### fluorescence

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#### Resonant photon excitation in EBIT



• Synchrotron radiation (PETRAIII),

• Free-electron lasers (LCLS), provide X-rays with high power and energy resolution



Soft X-ray laser spectroscopy at FLASH FEL 50 eV beam excites resonantly the *2s-2p* transition in Li-like Fe<sup>23+</sup> at 50 eV



#### S. W. Epp et al., Phys. Rev. Lett. 98 (2007) 183001



## LCLS: Linac Coherent Light Source



Photon bunches:
2 mJ
10...300 fs
120 Hz
550 eV...11000 eV
EBIT at SXR



last km

of SLA

300 m of

undulators

#### Going into the X-ray region at SXR (LCLS)

#### Soft X-ray laser (up to 2000 eV) monochromatic photon beam

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CAUTIO





## Adding X-ray fluorescence detectors





#### New results: Overview spectra of Br<sup>33+</sup> (Li-like)



#### S. Bernitt, MPIK (2016)

## High resolution scans show line widths



#### X-ray energy scanned by monochromator

#### Overview and detailed scans

FÜR KERNPHYSI



#### Linewidth determinations in Fe<sup>23+</sup> and Fe<sup>24+</sup>

			MCDF			
Line	Experiment	Theory	Radiative	Auger	Total	
W	311 (10)	301 [49]	301	0	301	
		301 [35]				
		303 [33]				
		315 [50]				
q	255 (31)	310 [33]	312	0.081	312	
r	250 (11)	226 [33]	208	32	241	
t	131 (29)	167 [33]	112	52	163	
E1	437 (12)	382 [33]	292	95	387	
E2	178 (34)	149 [33]	31	136	167	
<i>C</i> 1	524 (12)	499 [33]	218	334	552	
<i>C</i> 2	385 (207)	611 [33]	412	135	547	
N1	565 (45)	498 [33]	146	400	546	
N2	594 (50)	504 [33]	164	338	502	
<i>N</i> 3	570 (109)	505 [33]	155	360	515	
01	859 (229)	756 [33]	264	529	793	
02	772 (228)	785 [33]	304	518	822	
F	998 (203)	989 [33]	351	651	1020	

Natural linewidths wL of  $1s \rightarrow 2p$  transitions (in meV) J. K. Rudolph et al., PRL 111, 103002 (2013)

## Transition energies

TABLE I. X-ray transitions of heliumlike to fluorinelike iron ions resonantly excited from the ground state with synchrotron radiation. X-ray fluorescence was detected as a function of photon energy. Energies are given in units of eV. The calibration is based on the absorption edge technique. The experimental uncertainties are shown as (statistical)(systematic). Relative energies are not affected by the systematic uncertainty which accounts for a shift of the absolute scale. Angle brackets enclose results affected in their accuracy by line blends.

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Ion	Line	Initial state	Final state	This experiment	Theory		Theory		Experiment	
Fe <sup>24+</sup>	w	$1s^{2} {}^{1}S_{0}$	$1s 2p P_1$	6700.549 (5) (70)	6700.4347 (11)	[32]	6700.4	[33]	6700.8	[22]
					6700.490	[34]	6700.4	[35]	6700.4	[23]
									6700.9	[36]
$Fe^{24+}$	У	$1s^{2} {}^{1}S_{0}$	$1s 2p {}^{3}P_{1}$	6667.671 (3) (69)	6667.5786 (12)	[32]	6667.6	[33]	6667.9 (4)	[22]
					6667.629	[34]	6667.6	[35]	6667.5	[23]
22.1			2						6667.5	[36]
$\mathrm{Fe}^{23+}$	t	$1s^2 2s  {}^2S_{1/2}$	$1s 2s 2p^2 P_{1/2}$	6676.202 (3) (69)	6676.129 (47)	[37]	6676.4	[33]	$\langle 6676.8(7) \rangle$	[22]
22.1					6675.8	[38]			6676.3	[23]
Fe <sup>23+</sup>	q	$1s^2 2s  {}^2S_{1/2}$	$1s 2s 2p^2 P_{3/2}$	6662.240 (6) (69)	6662.188 (11)	[37]	6661.9	[33]	6662.1 (5)	[22]
22.1					6661.9	[38]			6662.2	[23]
Fe <sup>23+</sup>	r	$1s^2 2s  {}^2S_{1/2}$	$1s 2s 2p^2 P_{1/2}$	6652.826 (3) (69)	6652.776 (25)	[37]	6653.5	[33]	(6654.2(7))	[22]
22.1					6652.6	[38]			6652.5	[23]
$\mathrm{Fe}^{23+}$	и	$1s^2 2s  {}^2S_{1/2}$	$1s 2s 2p {}^{4}P_{3/2}$	6616.629 (4) (68)	6616.559 (11)	[37]	6616.7	[33]	(6617.9(1.2))	[22]
22.1		2 21	2 1						6616.6	[23]
Fe <sup>22+</sup>	E1	$1s^2 2s^2 {}^1S_0$	$1s 2s^2 2p P_1$	6628.804 (5) (68)	6631.057	[39]	6628.7	[33]	6628.9 (3)	[22]
		2 - 2 1 -			6627.4 <sup>a</sup> /6628.3 <sup>b</sup>		6627.39	[40]	6628.7	[23]
Fe <sup>22+</sup>	E2	$1s^2 2s^2 {}^1S_0$	$1s2s^22p^3P_1$	6597.858 (3) (67)	6596.55	[40]	6595.8	[33]		
- 21 /					6596.1°/6597.7°					
Fe <sup>21+</sup>	В	$1s^2 2s^2 2p^2 P_{1/2}$	$1s 2s^2 2p^2 {}^2P_{1/2}$	(6586.085 (7) (67))	6586.3ª/6585.1°		6586.3	[33]	(6585.9(5))	[22]
			$1s 2s^2 2p^2 {}^2D_{3/2}$	(6586.085 (7) (67))	6587.0°/6585.8°		6586.5	[33]	(6585.7)	[23]
- 20	~ .				(6587.2)	[41]				
Fe <sup>20+</sup>	C1	$1s^2 2s^2 2p^2 {}^3P_0$	$1s 2s^2 2p^3 {}^3D_1$	6544.225 (4) (66)	6544.8ª/6544.0°		6543.6	[33]	(6544.6(9))	[22]
- 20									(6544.4)	[23]
$Fe^{20+}$	<i>C</i> 2	$1s^2 2s^2 2p^2 {}^{3}P_0$	$1s 2s^2 2p^3 {}^{3}S_1$	6556.879 (16) (66)	6557.3°/6556.3°		6555.0	[33]		
Feigh	NI	$1s^2 2s^2 2p^3 + S_{3/2}$	$1s 2s^2 2p^{44}P_{5/2}$	6497.067 (5) (65)	6497.5*/6497.2*		6496.6	[33]	(6497.7(1.4))	[22]
₽ 10±		1 2 2 2 2 3 4 9	1 2 2 2 4 2 5		crom payers of oh		65060		6497.3	[23]
Ferr	N2	$1s^2 2s^2 2p^{34}S_{3/2}$	$1s 2s^2 2p^{-2}P_{3/2}$	6506.845 (7) (65)	6507.3%/6506.9°		6506.0	[33]	(6509.6(1.4))	[22]
<b>D</b> 10+	110	1 2 2 2 2 340	1 2 2 2 4 4 5	(500 100 (14) (65)	crop caucrop th		6500.1	52.23	(6509.1)	[23]
Ferr	N3	$1s^2 2s^2 2p^{34}S_{3/2}$	$1s 2s^2 2p^{-1/2}$	6509.133 (14) (65)	6509.6*/6509.1*		6508.1	[33]	(6509.6(1.4))	[22]
r 18+	01	1 2 2 2 2 4 3 0	1 2 2 2 5 3 2		CACT ABICACC CD		(5 ( A A	5223	(6509.1)	[23]
Ferei	01	$1s^2 2s^2 2p^4 {}^{5}P_2$	$1s 2s^2 2p^3 P_2$	6466.900 (14) (64)	6467.476466.5	F 4 13	6564.4	[33]	6467.6 (1.7)	[22]
E-18+	02	1 -2 2 -2 2 -4 3 D	1 · 2 · 2 · 5 3 D	6474219 (22) (64)	0400.3	[41]	6472.0	[22]	(0400.3)	[23]
rent	02	$1s^2 2s^2 2p^{+3}P_2$	$1s 2s^2 2p^3 P_1$	04/4.318 (33) (64)	64727	F417	04/3.0	[33]	(04/2./(2./))	[22]
E-17+	Б	1-20-20-520	1 - 2 - 2 2 - 6 2 6	(425.220 (14) (62)	04/3.7	[41]	6424.9	[22]	(04/4.7)	[23]
Fe <sup>rr</sup>	F	$1s^{-}2s^{-}2p^{-}P_{3/2}$	$1s 2s^2 2p^{\circ 2}S_{1/2}$	6435.239 (14) (63)	0435./~/0434.6°		0434.8	[55]	6436.1 (2.0)	[22]
									6434.8	23

<sup>a</sup>Our theoretical results obtained in the framework of the multiconfiguration Dirac-Fock (MCDF) method [29] <sup>b</sup>Our theoretical results using the Flexible Atomic Code (FAC) of Gu with the standard configuration-interaction package [42]. 16 dominant transitions measured with high precision

J. K. Rudolph et al., PRL 111, 103002 (2013)





#### Petra III: Kr<sup>34+</sup> resonant excitation



Experimental raw data (Si(333)) of the w and y line along with the Pb L3edge absorption spectrum (PbL) for energy calibration. The w data points show a result from a single scan (15 minutes). Fitted Bragg angle is 26.89274(4) degrees. The y line is an overlay of 25 single scans. PbL (open circles): our data scaled to the (dashed) reference data . Inset (blue): Centroids from12 individual scans of w.

Epp et al, PRA 92, 020502(R) (2015)



#### Fluorescence vs. photoion yield



#### R. Steinbrügge et al., PRA **91**, 032502 (2015)



## Transition energies for He-like and Li-like ions

- 70 meV systematic uncertainty of the calibration
- 5 meV typical statistical uncertainties
- For Z>18 never measured so precisely

#### J. K. Rudolph et al., PRL 111, 103002 (2013)



#### PolarX-EBIT and gas-cell setup





#### Some theory models vs. experiments





#### O<sub>2</sub> photoabsorption



From: Feifel et al.



#### He-like $O^{5+}$ 1s-2p (line q)





#### O<sub>2</sub> absorption edges




#### Overview spectrum at low resolution and statistics





#### ≈500 meV offset relative to old calibration!



G.V. Brown, P. Beiersdorfer, T. Lockard (LLNL); R. Kelley, C.A. Kilbourne, M.A. Leutenegger, F.S. Porter (GSFC); J. Wilms (FAU)

Lawrence Livermore National Laboratory





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#### Photon energy range accessed







#### Photoionization studies





### Ar<sup>12+</sup>: ionization potential 684 eV



**Figure 4.** (color online) PI of ions in charge state q=12 demonstrate the ability to access HCIs by the EBIT based method. (a) PI threshold of Ar<sup>12+</sup> at 683.93 eV, obtained by a fit to the Ar<sup>13+</sup>-signal. The strong background signal below the ionization edge is mainly caused by residual gas ions with close lying charge-to-mass ratios (O<sup>5+</sup>, C<sup>4+</sup>). (b) Near-threshold PI of Fe<sup>12+</sup> revealing two significant resonances at 362.6 eV and 370.0 eV.

## Fe<sup>14+</sup> photoionization





### Fe<sup>14+</sup> photoionization



Strong resonances allow high resolution nearly reaching natural line width Current sensitivity for non-resonant photoionization around 20 kbarn



# Electron-driven resonant processes





- Fe photorecombination studies
- Slow scans at high electron energy resolution





# High statistics





- Fast Fe photorecombination studies
- Maxwellian electron energy distribution





#### lons in any desired charge state can be prepared, stored and spectroscopically studied

#### T. M. Bauman et al., PRA **90**, 052704 (2014)

#### Unexpected, strong contributions by many-electron resonant excitation at high resolution



C. Shah et al., Phys. Rev. E 93, 061201(R) (2016)
C. Beilmann et al., Phys. Rev. Lett 107, 143201 (2011)
C. Beilmann et al., Phys. Rev. A 88, 062706 (2013)

#### Unexpected, strong contributions by many-electron resonant excitation at high resolution



C. Shah et al., Phys. Rev. E 93, 061201(R) (2016)
C. Beilmann et al., Phys. Rev. Lett 107, 143201 (2011)
C. Beilmann et al., Phys. Rev. A 88, 062706 (2013)



#### L-shell resonant excitation





#### Mean charge state of plasmas changed by many-electron resonant excitation



C. Shah et al., Phys. Rev. E 93, 061201(R) (2016)

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### Fe L-shell resonant excitation



#### Pedro Amaro, in preparation

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# Conclusions from FEL, synchrotron and photorecombination studies

- Previously known but neglected complex multielectron processes have experimentally shown unexpectedly strong contributions to both photrecombination and photoionization
- Inclusion of those channels in calculations is necessary to achieve agreement with existing experimental data
- Even the most advanced theoretical methods are not as accurate as the experiments
- Achieving agreement at the 1% (energies) respectively 5% (cross sections) levels will require more dedicated theoretical work and benchmarking experimental data



## Optical and EUV spectroscopy with EBITs



### Spectroscopy of few-electron ions in the visible range



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#### Fe<sup>13+</sup> (Fe XIV): the *"green coronal line"*



Hendrik Bekker, accepted PRA 2018



#### Level crossings at $Ir^{17+}$ provide $\alpha$ sensitivity



- At certain charge state levels change order
- 4f goes below 5s in  $Ir^{17+}$
- Opposite parities degenerate: 4f<sup>12</sup> 5s<sup>2</sup>, 4f<sup>13</sup> 5s, 4f<sup>14</sup>
- Many slow M1, E1, E2, M2, M3 transitions become possible
- Several long lived "ground states" available

Windberger et al., PRL 114, 150801 (2015)



#### Visible spectra of M1 lines in Ir ions



#### First observations; line identification difficult



### Improved resolving power





Transition	Wavelength (nm)	Updated
${}^{3}P_{1} - {}^{3}F_{2}$	226.63(2)	
${}^{3}F_{3} - {}^{3}F_{3}{}^{o}$	240.098(1)	
${}^{1}G_{4} - {}^{3}F_{3}{}^{o}$	255.8684(3)	
${}^{3}H_{4} - {}^{3}H_{5}$	324.6113(4)	
${}^{1}F_{3}^{o} - {}^{3}F_{4}^{o}$	329.3025(4)	329.3028 (2)
${}^{1}D_{2} - {}^{3}F_{2}$	365.0318(10)	365.03230(14)
${}^{1}F_{3}{}^{o} - {}^{3}F_{3}{}^{o}$	390.7408(9)	
${}^{3}F_{3} - {}^{3}F_{4}$	391.8220(3)	391.82393(16)
${}^{3}P_{2} - {}^{1}D_{2}$	399.4702(3)	
${}^{3}H_{5} - {}^{3}H_{6}$	422.8950(3)	422.89512(17)
${}^{1}D_{2} - {}^{3}F_{3}$	431.6044(3)	431.60376(8)
${}^{1}G_{4} - {}^{3}F_{4}$	435.6348(5)	435.7595 (2)
${}^{3}H_{4} - {}^{1}G_{4}$	445.7057(9)	445.70825(17)
${}^{3}F_{2}{}^{o} - {}^{3}F_{3}{}^{o}$	482.7039(7)	482.70386(8)
${}^{3}H_{4} - {}^{3}F_{3}$	503.423(2)	
${}^{3}P_{1} - {}^{1}D_{2}$	597.65(2)	

#### H. Bekker et al., in preparation

J. R. Crespo López-Urrutia, MPIK: Charge-state resolving analysis, EUV-Soft X-Ray Sources Workshop 2016



#### Comparison between theories



- Fock-space coupled cluster calculation (A. Borschevski) shows agreement with experimental result at a level suitable for identification.
- Its deviations from experiment are smaller than the average separation between spectral lines (as given by the green band).
- \* Berengut *et al.*, PRL **106**, (2011) Windberger et al., PRL **114**, 150801 (2015)

J. R. Crespo López-Urrutia, MPIK: Charge-state resolving analysis, EUV-Soft X-Ray Sources Workshop 2016



#### Pm-like (61 electrons) isoelectronic sequence



Wavelength (nm)

Isoelectronic sequence studied in detail to find analogies



## Electron beam ion trap diagnostics



#### EUV grating spectrometer 2



Si drift detector, Ge detector X-ray photons 1 to 30 keV Two grating spectrometers EUV photons 40 eV to 1 keV Metallic magnetic microcalorimeter X-ray photons 2 to 8 keV



#### Hinode ("sunrise"): space telescope (2006) studying the solar corona



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#### Studies of Fe HCI with charge-state resolution



Gu et al., Astrophys. J. 696, 2275 (2009)



# Understanding optical and EUV spectra of Sn ions



#### Understanding Sn spectra





#### Understanding Sn spectra



#### F. Torretti, ARCNL, PRA (2017)



#### Understanding Sn spectra



#### A. Windberger et al., Phys. Rev A 94, 012506 (2016)


## Understanding Sn spectra



#### F. Torretti, ARCNL, PRA (2017)



## Understanding Sn spectra



#### A. Windberger et al., Phys. Rev A 94, 012506 (2016)



## Understanding Sn spectra



#### F. Torretti, ARCNL, PRA (2017)



## Optical frequnecy metrology with HCI stabilization

## Production, deceleration, implantation of HCI



Lisa Schmöger et al., Science **347**, 1233 (2015)



## HCI identification by image analysis



- The single HCI (here Ar<sup>13+</sup>) repels Be<sup>+</sup> ions and produces a hole in the Coulomb crystal
- Addressing a single ion in the trap with a focused beam is possible due to large separation.
   Lisa Schmöger et al., Science 347, 1233 (2015)

## Setup at PTB





## Next generation cryogenic trap



- Cryogenic, XUHV
- Ultra-low vibration
- Superconducting high-Q RF resonator



## Table-top EBITs for PTB, Petra-III, Blaum division





#### Peter Micke et al., RSI (2018)



- Whole new class of laser-accessible targets, with Z and ionic charge as parameters
- Great variety of optical and EUV lines from fine and hyperfine transitions up to the highest charge states
- Stable up to X-ray region
- Forbidden transitions suitable as frequency standards
- Low sensitivity to DC, AC Stark, Zeeman and blackbody shifts
- Highest sensitivity to fine-structure constant α in atomic systems



## A laser high harmonic frequency comb in the VUV



## Temperature-controlled container for HHG-frequency comb





# Charge exchange studies with photon emission



## Charge exchange from neutrals to HCI



Solar wind: Beiersdorfer, P., Boyce, K. R., Brown, G. V., et al. 2003, Science, **300**, 1558 (2003) Beiersdorfer, P., Schweikhard, L., Liebisch, P., & Brown, G. V., ApJ, **672**, 726 (2008) Galaxy Clusters: L. Gu, ... C. Shah,... et al, A&A **611**, A26 (2018)



## Charge exchange in the laboratory



#### Shah, C., Dobrodey, S., Bernitt, S., JRCLU et al., ApJ 833, 52 (2016)





#### Shah, C., Dobrodey, S., Bernitt, S., JRCLU et al., ApJ 833, 52 (2016)





#### Shah, C., Dobrodey, S., Bernitt, S., JRCLU et al., ApJ 833, 52 (2016)





Shah, C., Dobrodey, S., Bernitt, S., JRCLU et al., ApJ 833, 52 (2016)





Shah, C., Dobrodey, S., Bernitt, S., JRCLU et al., ApJ 833, 52 (2016)

#### L. Gu, ... C. Shah, ... et al, A&A 611, A26 (2018)



## Summary

- X-ray Lyman- $\alpha$  and He-like studies with ppm accuracy
- EUV, VUV studies with ppm accuracy
- Photoionization studies from N<sup>3+</sup> to Fe<sup>23+</sup>
- X-ray laser spectroscopy demonstrated:
  - $Fe^{15+...17+}$  resonance transitions at ~800 eV
  - Ly-series lines of  $F^{8+}$  and  $O^{7+}$  at ~800 eV
  - Fe K $\alpha$  in Fe<sup>21+...24+</sup> ions excited at 6.7 keV
  - Kr<sup>34+</sup>, Br<sup>33+</sup> high-resolution studies up to 14 keV
  - Soft x-ray calibrations with  $\Delta E \approx 20 \text{ meV}$
- Oscillator strengths, line widths, cross sections determinations
- Charge exchange with photon emission
- VUV frequency comb in preparation
- Optical clock tests at PTB already running in collaboration



## MPIK current and former team members

T. M. Baumann, C. Beilmann, H. Bekker, S. Bernitt,
M. Blessenohl, S. Bogen, A. Borodin, G. Brenner,
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