



Atomic structure and radiative data calculations for heavy elements of interest in fusion plasma research

Theoretical challenges and recent advances



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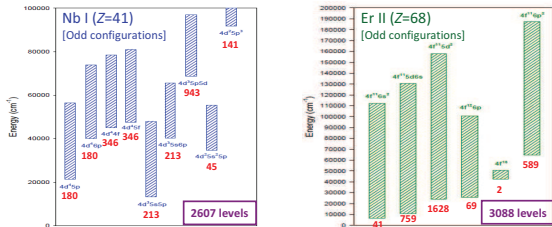
Introduction – Motivations

The Periodic Table of the Elements

1																	2
H																	He
3	4															10	
Li	Be															Ne	
6.941	9.012182															20.1797	
11	12															18	
Na	Mg															Ar	
22.989769	24.304															39.962383	
13	14	15	16	17	18											36	
Al	Si	P	S	Cl	Ar												
26.981538	28.08558	30.973762	32.06	35.45	39.962383												
5	6	7	8	9	10											18	
B	C	N	O	F	Ne												
10.811	12.0107	14.00643	15.999	18.998403	20.1797												
21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36		
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se		
39.0983	40.078	44.955912	47.88	50.9415	51.9961	54.938045	55.845	58.933195	58.6934	63.546	65.38	69.723	72.630	74.9216	78.96		
37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52		
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te		
85.4678	87.62	88.90584	91.224	92.90638	95.94	98.90625	101.07	102.9055	106.90558	107.8682	112.4118	114.818	118.710	121.757	127.603		
55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70		
Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po		
132.90545	137.327	138.90471	178.49	180.94788	183.84	186.207	190.23	192.222	195.083	196.96657	200.59	204.3833	207.2	208.9804	209		
87	88	89	90	91	92	93	94	95	96	97	98	99	100	101	102		
Fr	Ra	Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No		
223	226	227	232.0377	231.03688	238.02891	237.04817	244.0642	243.06138	247.07035	247.07035	251.07958	252.083	257.10528	258.10	259.10874		
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Introduction – Motivations

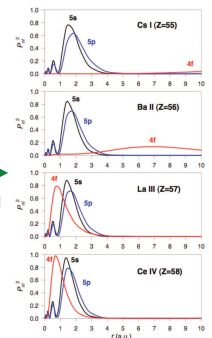
Fifth row
Sixth row
Lanthanides



Introduction – Motivations

Fifth row
Sixth row
Lanthanides

Collapse of the 4f orbital



Introduction – Motivations

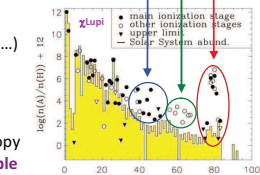
Fifth row
Sixth row
Lanthanides

Very complex spectroscopic structures (unfilled 4d, 5d, 4f shells)
→ little investigated up to now (both theoretically and experimentally)

Interest in many other fields :

- Astrophysics (CP stars, nucleosynthesis, ...)
- Solid state physics (doped crystals, ...)
- New light sources (lasers, lamps, ...)
- Plasma physics (fusion, ITER, ...)

Progress in computers & laser spectroscopy
→ New systematic investigations possible



Determination of atomic data

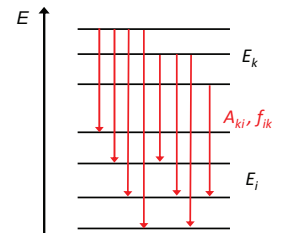
Theoretically and experimentally...

Atomic structure parameters

- Energy levels, E_k, E_i
- Transition energies, ΔE_{ki}
- Wavelengths, λ_{ki}

Radiative parameters

- Transition probabilities, A_{ki}
- Oscillator strengths, f_{ik}
- Radiative lifetimes, $\tau_k = 1/\sum_i A_{ki}$



Theoretical approach

Relativistic Hartree-Fock method (HFR)

(R.D. Cowan, *The Theory of Atomic Structure and Spectra*, Univ. California Press, Berkeley, 1981)

Based on the Schrödinger equation (atom with N electrons)

$$H\Psi = E\Psi \quad \text{with} \quad H = \sum_{i=1}^N \left[-\frac{\hbar^2}{2m} \nabla_i^2 - \frac{Ze^2}{4\pi\epsilon_0 r_i} + \sum_{j<i} \frac{e^2}{4\pi\epsilon_0 r_{ij}} \right]$$

Central field approximation

$$\varphi_{nlm_l m_s}(\mathbf{q}) = \frac{1}{r} P_{nl}(r) Y_{lm_l}(\theta, \phi) \chi_{1/2, m_s}(\sigma)$$

Slater determinant

$$\Psi(\mathbf{q}_1, \mathbf{q}_2, \dots, \mathbf{q}_N) = \frac{1}{\sqrt{N!}} \begin{vmatrix} \varphi_1(\mathbf{q}_1) & \varphi_2(\mathbf{q}_1) & \dots & \varphi_N(\mathbf{q}_1) \\ \varphi_1(\mathbf{q}_2) & \varphi_2(\mathbf{q}_2) & \dots & \varphi_N(\mathbf{q}_2) \\ \dots & \dots & \dots & \dots \\ \varphi_1(\mathbf{q}_N) & \varphi_2(\mathbf{q}_N) & \dots & \varphi_N(\mathbf{q}_N) \end{vmatrix}$$

Theoretical approach

Relativistic Hartree-Fock method (HFR)

(R.D. Cowan, *The Theory of Atomic Structure and Spectra*, Univ. California Press, Berkeley, 1981)

Hartree-Fock equations

$$\left[-\frac{\hbar^2}{2m} \frac{d^2}{dr^2} + \frac{\hbar^2 l(l+1)}{2mr^2} - \frac{Ze^2}{4\pi\epsilon_0 r} \right] P_i(r) + \left[\sum_{j=1}^N \int |P_j(r)|^2 \frac{e^2}{4\pi\epsilon_0 r_{ij}} d\tau_j \right] P_i(r) - \left[\sum_{j=1}^N \delta_{m_i m_j} \int P_j^*(r) P_i(r) \frac{e^2}{4\pi\epsilon_0 r_{ij}} d\tau_j \right] P_i(r) = E_i P_i(r)$$

Resolution of Hartree-Fock equations

Iterative method → self-consistent field method

Theoretical approach

Relativistic Hartree-Fock method (HFR)

(R.D. Cowan, *The Theory of Atomic Structure and Spectra*, Univ. California Press, Berkeley, 1981)

Multiconfiguration approach

$$\Psi^k = \sum_b a_b^k \Psi_b$$

Relativistic effects

Included perturbationally (spin-orbit, mass-velocity, Darwin term)

Good agreement with fully relativistic methods

Ab initio or semi-empirical approach

Experimental energy levels can be used to optimize the radial parameters

Theoretical approach

Relativistic Hartree-Fock method (HFR)

(R.D. Cowan, *The Theory of Atomic Structure and Spectra*, Univ. California Press, Berkeley, 1981)

Multiconfiguration approach

$$\Psi^k = \sum_b a_b^k \Psi_b$$

Example : Lu III (Lu^{2+})

(Ground config. : $[\text{Xe}]4f^{14}5s^25p^66s$)

Intravalance correlation
(up to $n=6$, $l=3$)

Even parity	Odd parity
$5s^25p^66s$	$5s^25p^66p$
$5s^25p^65d$	$5s^25p^65f$
$5s^25p^66d$	$5s^25p^66f$
5 levels	6 levels

Core-valence correlation (with $5s, 5p$)

Even parity	Odd parity	Even parity	Odd parity
$5s^25p^5d5f$	$5s^25p^5d^2$	$5s^25p^5d^2$	$5s^25p^5d5f$
$5s^25p^5d6p$	$5s^25p^5f^2$	$5s^25p^5f^2$	$5s^25p^5d6p$
$5s^25p^5d6f$	$5s^25p^6s^2$	$5s^25p^6s^2$	$5s^25p^5d6f$
$5s^25p^5f6s$	$5s^25p^6p^2$	$5s^25p^6p^2$	$5s^25p^5f6s$
$5s^25p^5f6d$	$5s^25p^6d^2$	$5s^25p^6d^2$	$5s^25p^5f6d$
$5s^25p^5f6p$	$5s^25p^6f^2$	$5s^25p^6f^2$	$5s^25p^5f6p$
$5s^25p^6s6f$	$5s^25p^5d6s$	$5s^25p^5d6s$	$5s^25p^6s6f$
$5s^25p^6p6d$	$5s^25p^5d6d$	$5s^25p^5d6d$	$5s^25p^6p6d$
$5s^25p^6d6f$	$5s^25p^5f6p$	$5s^25p^5f6p$	$5s^25p^6d6f$
	$5s^25p^6s6d$	$5s^25p^6s6d$	
	$5s^25p^6p6f$	$5s^25p^6p6f$	
	885 levels		904 levels

Core-polarization effects (HFR + CPOL)

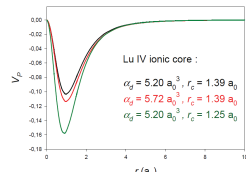
(see e.g. Quinet *et al.*, M.N.R.A.S. **307**, 934, 1999; Quinet *et al.*, J. Alloys Compd **344**, 255, 2002)

Intravalance correlation considered within a configuration interaction scheme

Core-valence correlation represented by a **core-polarization model potential** depending on two parameters (dipole polarizability α_d and cut-off radius r_c)

$$V_{P1} = -\frac{1}{2} \alpha_d \sum_{i=1}^n \frac{r_i^2}{(r_i^2 + r_c^2)^3}$$

$$V_{P2} = -\alpha_d \sum_{i>j} \frac{\vec{r}_i \cdot \vec{r}_j}{[(r_i^2 + r_c^2)(r_j^2 + r_c^2)]^{3/2}}$$



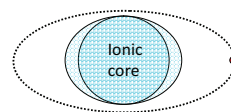
Correction to the dipole radial integral

$$\int_0^\infty P_{nl}(r) r P_{n'l'}(r) dr \quad \text{replaced by} \quad \int_0^\infty P_{nl}(r) r \left[1 - \frac{\alpha_d}{(r^2 + r_c^2)^{3/2}} \right] P_{n'l'}(r) dr$$

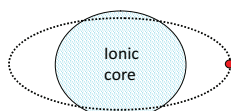
Core-polarization effects (HFR + CPOL)

(see e.g. Quinet *et al.*, M.N.R.A.S. **307**, 934, 1999; Quinet *et al.*, J. Alloys Compd **344**, 255, 2002)

Core-polarization



Core-penetration



Semi-empirical optimization

(R.D. Cowan, *The Theory of Atomic Structure and Spectra*, Univ. California Press, Berkeley, 1981)

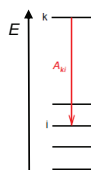
Radial parameters (average energies, electrostatic integrals, spin-orbit parameters) adjusted in order to minimize the discrepancies between computed Hamiltonian eigenvalues and experimental energy levels

→ Optimization of the transition energies and wavefunctions

→ Optimization of the radiative decay rates

$$A_{ki} = \frac{64\pi^4 e^4 a_0^2 (\Delta E_{ki})^3}{3h 2J_k + 1} \left| \left\langle \gamma, J_k \left\| P^{(1)} \right\| \gamma_i, J_i \right\rangle \right|^2$$

$$= 2.0261 \times 10^{-6} \frac{(\Delta E_{ki})^3}{2J_k + 1} \left| \left\langle \gamma, J_k \left\| P^{(1)} \right\| \gamma_i, J_i \right\rangle \right|^2$$



→ A good experimental knowledge of the level structure for the atom (or ion) considered is needed

A specific example (Lu III)

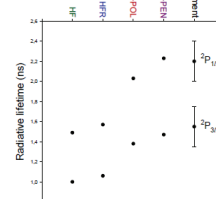
(Blémont *et al.*, J. Phys. B **32**, 3409, 1999)

Radiative lifetimes for $6p \ ^2P_{1/2}$ (38401 cm^{-1}) and $6p \ ^2P_{3/2}$ (44705 cm^{-1})

$1s^22s^22p^63s^23p^63d^{10}4s^24p^64d^{10}4f^{14}5s^25p^6 nl$

Lu³⁺ ionic core with 68 electrons :
 $\alpha_d = 5.20 \text{ a}_0^3$ (Fraga *et al.*, 1976)
 $r_c = \langle r \rangle_{5p} = 1.39 \text{ a}_0$

Method	$2P_{1/2}$	$2P_{3/2}$
HFR	1.49	1.00
HFR	1.57	1.06
HFR+POL	2.03	1.38
HFR+POL+PEN	2.23	1.47
Experiment	2.20 ± 0.20	1.55 ± 0.20



Experimental measurements

Time-resolved laser-induced fluorescence (TR-LIF)

(see e.g. Bergström *et al.*, Z. Phys. D **8**, 17, 1988; Xu *et al.*, Phys. Rev. A **70**, 042508, 2004)

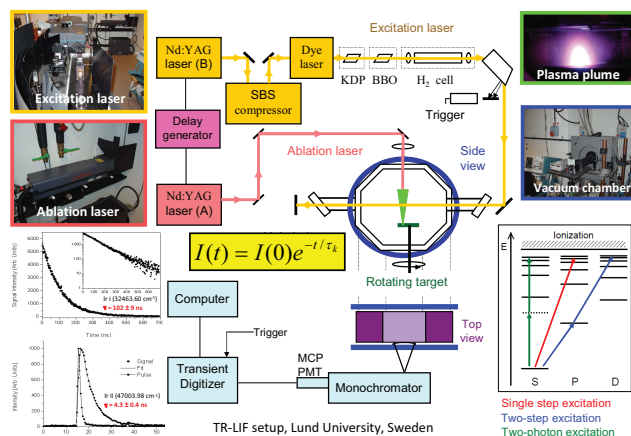
Accurate measurements of radiative lifetimes (within a few %)

Lifetime range from 1 ns to 300 ns

Selective excitation (no cascading problems)

Many levels accessible (using different laser dyes)

Different ionization degrees accessible in the laser-produced plasma (neutral, singly-, doubly- and trebly-ionized atoms)



Results obtained so far

Calculations and measurements performed during the period 1998 – 2014

Fifth row										Sixth row									
30 ions ~200 lifetimes										29 ions ~200 lifetimes									
1 H	4 Be	5 B	6 C	7 N	8 O	9 F	10 Ne	11 Na	12 Mg	13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	19 K	20 Ca	21 Sc	22 Ti
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr	37 Rb	38 Sr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe	55 Cs	56 Ba
55 Cs	56 Ba	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	72 Hf	73 Ta	74 W
73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn	87 Fr	88 Ra	89 Ac	90 Th	91 Pa	92 U
92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Ts

Lanthanides
30 ions
~300 lifetimes

Radiative lifetimes, transition rates (lanthanides)

Z	Ion	References
58	Ce II	Palmeri et al., Phys. Scr. 61 , 323 (2000); Zhang et al., Phys. Scr. 63 , 122 (2001)
	Ce III	Blémont et al., MNRAS 336 , 1155 (2002)
	Ce IV	Zhang et al., Phys. Rev. Lett. 87 , 27301 (2001)
59	Pr I	Blémont et al., J. Phys. B 37 , 1381 (2004)
	Pr II	Blémont et al., Eur. Phys. J. D 27 , 33 (2003)
	Pr III	Palmeri et al., ApJS 129 , 367 (2000); Blémont et al., Phys. Rev. A 64 , 022503 (2001)
	Pr IV	Enzonga Yoca & Quinet, J. Phys. B 46 , 145003 (2013)
60	Nd I	Blémont et al., J. Phys. B 37 , 1381 (2004)
	Nd II	Xu et al., MNRAS 346 , 433 (2003)
	Nd III	Zhang et al., A&A 385 , 724 (2002)
	Nd IV	Enzonga Yoca & Quinet, J. Phys. B 47 , 035002 (2014)
61	Pm II	Fivet et al., MNRAS 380 , 771 (2007)
62	Sm II	Xu et al., J. Phys. B 36 , 4773 (2003)
	Sm III	Blémont et al., A&A 399 , 343 (2003)
63	Eu III	Quinet & Blémont, MNRAS 340 , 463 (2003)
64	Gd III	Blémont et al., A&A 393 , 717 (2002)
65	Tb III	Blémont et al., Phys. Rev. A 65 , 052502 (2002)
66	Dy III	Zhang et al., MNRAS 334 , 1 (2002)
67	Ho III	Blémont et al., MNRAS 328 , 1085 (2001); Zhang et al., A&A 384 , 364 (2002)
68	Er II	Xu et al., J. Phys. B 36 , 1771 (2003)
	Er III	Blémont et al., MNRAS 321 , 481 (2001)
69	Tm II	Quinet et al., JQSRT 62 , 625 (1999)
	Tm III	Li et al., J. Phys. B 34 , 1349 (2001)
70	Yb II	Blémont et al., J. Phys. B 31 , 3321 (1998); Li et al., J. Phys. B 32 , 1731 (1999); Blémont et al., J. Phys. B 35 , 4743 (2002)
	Yb III	Blémont et al., J. Phys. B 34 , 1869 (2001); Zhang et al., Eur. Phys. J. D 15 , 301 (2001)
	Yb IV	Wyart et al., Phys. Scr. 63 , 113 (2001)
71	Lu I	Fedchak et al., ApJ 542 , 1109 (2000); Dai et al., J. Phys. B 36 , 479 (2003)
	Lu II	Quinet et al., MNRAS 307 , 934 (1999); Fedchak et al., ApJ 542 , 1109 (2000)
	Lu III	Blémont et al., J. Phys. B 32 , 3409 (1999); Fedchak et al., ApJ 542 , 1109 (2000)

Radiative lifetimes, transition rates (fifth row)

Z	Ion	References
37	Rb III	Zhang et al., Eur. Phys. J. D 68 , 104 (2014)
39	Y II	Blémont et al., MNRAS 414 , 3350 (2011)
	Y III	Blémont et al., MNRAS 414 , 3350 (2011)
40	Zr I	Malcheva et al., MNRAS 395 , 1523 (2009)
	Zr II	Malcheva et al., MNRAS 367 , 754 (2006)
	Nb I	Nilsson et al., A&A 511 , A16 (2010)
	Nb II	Nilsson et al., A&A 511 , A16 (2010)
	Nb III	Nilsson et al., A&A 511 , A16 (2010)
42	Mo I	Quinet, J. Phys. B 35 , 19 (2002); Lundberg et al., J. Phys. B 43 , 085004 (2010)
	Mo III	Quinet, Phys. Scr., in press (2014)
	Tc I	Palmeri et al., MNRAS 363 , 452 (2005)
	Tc II	Palmeri et al., MNRAS 374 , 63 (2007)
	Ru I	Fivet et al., MNRAS 396 , 2124 (2009)
	Ru II	Palmeri et al., J. Phys. B 42 , 165005 (2009)
	Ru III	Palmeri et al., J. Phys. B 42 , 165005 (2009)
	Rh I	Quinet et al., J. Elec. Spectrosc. Rel. Phen. 184 , 174 (2011); Quinet et al., A&A 537 , A74 (2012)
	Rh III	Zhang et al., Phys. Scr. 88 , 065302 (2013)
46	Pd I	Xu et al., A&A 452 , 357 (2006)
	Pd II	Quinet, Phys. Scr. 54 , 483 (1996)
	Pd III	Zhang et al., Phys. Scr. 88 , 065302 (2013)
47	Ag II	Blémont et al., J. Elec. Spectr. Rel. Phen. 144-147 , 27 (2005); Campos et al., MNRAS 363 , 905 (2005)
48	Cd I	Zhang et al., Phys. Scr. 88 , 065302 (2013)
	Cd II	Blémont et al., Phys. Rev. A 62 , 032512 (2000); Xu et al., Phys. Rev. 70 , 042508 (2004)
49	In II	Xu et al., Phys. Rev. 70 , 042508 (2004)
	In III	Blémont et al., Phys. Rev. A 62 , 032512 (2000)
50	Sn I	Zhang et al., Phys. Rev. A 78 , 022505 (2008); Zhang et al., Eur. Phys. J. 55 , 1 (2009)
	Sn III	Blémont et al., Phys. Rev. A 62 , 032512 (2000)
51	Sb I	Hartman et al., Phys. Rev. A 82 , 052512 (2010)
52	Te II-III	Zhang et al., Astron. Astrophys. 551 , A136 (2013)

Radiative lifetimes, transition rates (sixth row)

Z	Ion	References
56	Ba I	Zhang et al., Phys. Rev. A 82 , 042507 (2010)
57	La I	Blémont et al., Eur. Phys. J. D 30 , 157 (2004)
	La II	Blémont et al., J. Phys. B 32 , 3409 (1999)
72	Hf I	Malcheva et al., MNRAS 396 , 2289 (2009)
	Hf III	Malcheva et al., MNRAS 396 , 2289 (2009)
73	Ta I	Fivet et al., Eur. Phys. J. D 37 , 29 (2006)
	Ta II	Quinet et al., A&A 493 , 711 (2009)
	Ta III	Fivet et al., J. Phys. B 41 , 015702 (2008)
74	W I	Quinet et al., J. Phys. B 44 , 145005 (2011)
	W II	Nilsson et al., Eur. Phys. J. D 49 , 13 (2008)
	W III	Palmeri et al., Phys. Scr. 78 , 015304 (2008)
	W IV	Enzonga Yoca et al., J. Phys. B 45 , 035001 (2012)
	W V	Enzonga Yoca et al., J. Phys. B 45 , 065001 (2012)
	W VI	Enzonga Yoca et al., J. Phys. B 45 , 035002 (2012)
75	Re I	Palmeri et al., Phys. Scr. 74 , 297 (2006)
	Re II	Palmeri et al., MNRAS 362 , 1348 (2005)
	Os I-II	Quinet et al., A&A 448 , 1207 (2006)
77	Ir I-II	Xu et al., JQSRT 104 , 52 (2007)
78	Pt II	Quinet et al., Phys. Rev. A 77 , 022501 (2008)
79	Au I	Fivet et al., J. Phys. B 39 , 3587 (2006); Blémont et al., MNRAS 380 , 1581 (2007)
	Au II	Fivet et al., J. Phys. B 39 , 3587 (2006); Blémont et al., MNRAS 380 , 1581 (2007)
	Au III	Enzonga Yoca et al., Phys. Scr. 78 , 025303 (2008)
80	Hg I	Blagoev et al., Phys. Rev. A 66 , 032509 (2002)
81	Tl I	Blémont et al., J. Phys. B 38 , 3547 (2005)
82	Pb I	Li et al., Phys. Rev. A 57 , 3443 (1998); Blémont et al., MNRAS 312 , 116 (2000)
	Pb II	Quinet et al., J. Phys. B 40 , 1705 (2007)
83	Bi I	Palmeri et al., Phys. Scr. 63 , 468 (2001)
	Bi III	Quinet et al., J. Phys. B 40 , 1705 (2007)

Some recent results : Nb II, Mo II, Rh II

Nb II : lifetimes measured for 17 energy levels belonging to 4d³5p

Level	E (cm ⁻¹)	τ (ns) [Theory] This work	τ (ns) [Experiment] This work	Previous
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	34886.35	5.9	5.6 ± 0.3	5.5 ± 0.3 ^a 5.7 ± 0.3 ^b
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	35520.82	5.7	5.5 ± 0.3	5.7 ± 0.3 ^a 5.7 ± 0.3 ^b
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	37298.24	2.7	2.6 ± 0.2	3.1 ± 0.3 ^a
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	37797.32	2.8	2.7 ± 0.2	3.2 ± 0.3 ^a
4d ⁴ (⁴ F)5p ⁴ G _{5/2}	40103.61	5.0	5.1 ± 0.3	5.2 ± 0.3 ^a
4d ⁴ (⁴ F)5p ⁴ P _{3/2}	41710.14	8.4	8.5 ± 0.7	
4d ⁴ (⁴ F)5p ⁴ P _{1/2}	42132.65	7.2	7.2 ± 0.5	
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	42596.55	7.2	7.5 ± 0.7	
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	43290.34	7.3	7.5 ± 0.4	
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	43618.39	6.7	6.7 ± 0.3	
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	43649.19	8.1	7.2 ± 0.4	12.8 ± 0.3 ^b
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	43887.08	6.3	5.7 ± 0.5	
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	44066.67	6.7	6.4 ± 0.5	5.3 ± 0.3 ^b
4d ⁴ (⁴ F)5p ⁴ P _{3/2}	44285.94	7.5	7.0 ± 0.5	
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	44924.60	6.6	6.2 ± 0.3	
4d ⁴ (⁴ F)5p ⁴ D _{5/2}	44970.71	5.8	5.2 ± 0.5	
4d ⁴ (⁴ F)5p ⁴ D _{3/2}	45802.47	4.6	4.6 ± 0.3	

^a Salih & Lawer (1983); ^b Hannaford et al. (1985) Mean ratio (Th/Exp) = 1.035 ± 0.048

Some recent results : Nb II, Mo II, Rh II

Mo II : lifetimes measured for 16 energy levels belonging to 4d⁴5p

Level	E (cm ⁻¹)	τ (ns) [Theory] This work	τ (ns) [Experiment] This work	Previous
4d ⁴ (⁴ D)5p z ⁴ P _{3/2}	48022.815	4.8	4.9 ± 0.2	4.8 ± 0.4 ^a 4.9 ± 0.3 ^b 4.3 ± 0.3 ^a 4.0 ± 0.3 ^b
4d ⁴ (⁴ D)5p z ⁴ D _{3/2}	50302.913	4.2	4.0 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ F _{3/2}	51373.170	4.7	4.2 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ F _{5/2}	51732.690	4.6	4.2 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ F _{7/2}	52217.587	4.4	4.0 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ F _{9/2}	52843.296	4.3	4.0 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ D _{3/2}	54238.949	5.5	5.1 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ D _{5/2}	54687.002	5.2	5.0 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ D _{7/2}	55216.203	4.9	4.6 ± 0.3	
4d ⁴ (⁴ D)5p z ⁴ D _{9/2}	55706.937	4.7	4.5 ± 0.3	
4d ⁴ (⁴ H)5p z ⁴ H _{3/2}	57892.289	5.2	4.9 ± 0.3	
4d ⁴ (⁴ H)5p z ⁴ H _{5/2}	58196.928	5.2	4.9 ± 0.3	
4d ⁴ (⁴ H)5p z ⁴ H _{7/2}	58761.258	5.0	4.8 ± 0.3	
4d ⁴ (⁴ H)5p z ⁴ H _{9/2}	59492.157	4.7	4.6 ± 0.3	
4d ⁴ (⁴ F)5p z ⁴ D _{3/2}	59841.060	8.5	7.6 ± 1.0	
4d ⁴ (⁴ F)5p z ⁴ D _{5/2}	60992.764	7.5	6.6 ± 0.5	

^a Hannaford & Lowe (1983); ^b Sikström et al. (2001) Mean ratio (Th/Exp) = 1.068 ± 0.039

Some recent results : Nb II, Mo II, Rh II

Rh II : lifetimes measured for 17 energy levels belonging to 4d⁷5p

Level	E (cm ⁻¹)	τ (ns) [Theory] This work	τ (ns) [Experiment] This work	Previous
4d ⁷ (⁴ F)5p ⁴ F _{7/2}	56547.3	4.4	3.9 ± 0.5	
4d ⁷ (⁴ F)5p ⁴ F _{9/2}	57020.8	4.0	3.9 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ F _{5/2}	58358.5	4.2	3.8 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ D _{3/2}	59161.5	3.9	3.3 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ F _{7/2}	59698.6	4.2	3.8 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ G _{5/2}	59702.4	3.1	3.0 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ G _{7/2}	59729.4	4.0	3.5 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ D _{5/2}	60448.4	3.9	3.4 ± 0.4	
4d ⁷ (⁴ F)5p ⁴ G _{5/2}	61173.1	3.5	3.3 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ D _{3/2}	61355.9	4.0	3.7 ± 0.5	
4d ⁷ (⁴ F)5p ⁴ G _{7/2}	61939.8	3.6	3.3 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ G _{5/2}	62194.4	3.5	3.2 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ G _{3/2}	62288.3	3.6	3.3 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ F _{7/2}	62326.1	2.8	2.4 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ F _{5/2}	63454.9	2.5	2.3 ± 0.3	
4d ⁷ (⁴ F)5p ⁴ G _{5/2}	63959.5	3.7	3.4 ± 0.2	
4d ⁷ (⁴ F)5p ⁴ G _{3/2}	65321.2	2.4	2.0 ± 0.2	

Mean ratio (Th/Exp) = 1.107 ± 0.047

The specific case of tungsten (W I, W II, W III)

Transition rates in W I, W II and W III

W I : 508 experimentally known energy levels within the $5d^6s^2$, $5d^6s$, $5d^6s7s$, $5d^6p$, $5d^6s6p$ and $5d^6s^26p$ configurations (Kramida & Shirai, J. Phys. Chem. Ref. Data **35**, 423, 2006; Wyart, J. Phys. B **43**, 074018, 2010)

W II : 263 experimentally known energy levels within the $5d^5$, $5d^46s$, $5d^36s^2$, $5d^4p$, $5d^36s6p$ and $5d^36s^26p$ configurations (Kramida & Shirai, J. Phys. Chem. Ref. Data **35**, 423, 2006)

W III : 235 experimentally known energy levels within the $5d^4$, $5d^36s$, $5d^26s^2$, $5d^3p$ and $5d^26s6p$ configurations (Iglesias *et al.*, J. Res. Natl Inst. Stand. Tech. **94**, 221, 1989)

Transition probabilities and oscillator strengths calculated for :

2525 W I lines in the range 223 – 1010 nm (Quinet *et al.*, J. Phys. B. **44**, 145005, 2011)

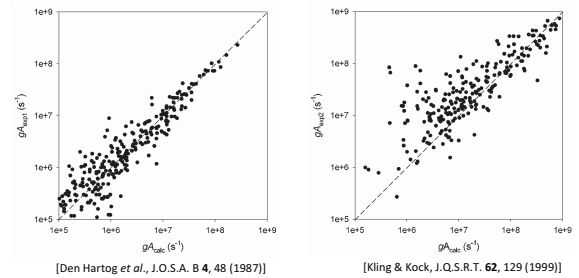
6086 W II lines in the range 143 – 990 nm (Nilsson *et al.*, Eur. Phys. J. D **49**, 13, 2008)

4822 W III lines in the range 83 – 1494 nm (Palmeri *et al.*, Phys. Scr. **78**, 015304, 2008)

The specific case of tungsten (W I, W II, W III)

Transition rates in W I

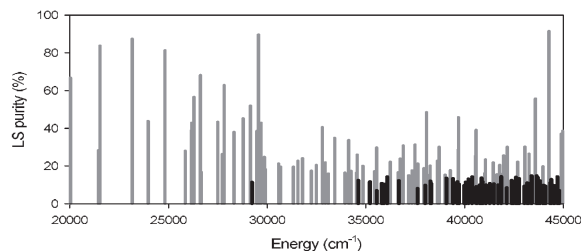
Comparison between theoretical transition probabilities and available experimental results



The specific case of tungsten (W I, W II, W III)

Transition rates in W I

Wavefunction purities (in LS coupling) for W I odd-levels below 45000 cm⁻¹ (bold lines represent purities smaller than 15%)



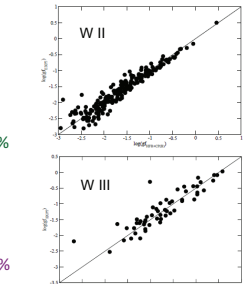
The specific case of tungsten (W I, W II, W III)

Transition rates in W II and W III

Comparison between theoretical and experimental lifetimes obtained in our work

E (cm ⁻¹)	J	Theory	Experiment
W II			
47179.941	3/2	6.8	5.8 ± 0.3
48284.498	5/2	7.6	5.9 ± 0.3
48181.034	9/2	20.4	21.5 ± 1.0
50430.999	3/2	5.3	6.6 ± 0.4
51045.292	7/2	4.3	4.7 ± 0.4
51254.429	3/2	2.2	2.1 ± 0.2
51438.064	5/2	4.7	4.6 ± 0.4
54498.608	7/2	2.9	2.1 ± 0.2
55392.446	9/2	2.4	2.3 ± 0.2
W III			
57231.04	2	2.9	2.9 ± 0.3
61488.36	3	2.3	2.5 ± 0.3

Comparison between theoretical oscillator strengths and available experimental results (W II : Kling *et al.*, JQSRT **67**, 227, 2000) (W III : Schultz-Johanning *et al.*, Phys. Scr. **63**, 367, 2001)



The specific case of tungsten (W I, W II, W III)

Critical evaluation of available data

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Radiative decay rates for W I, W II and W III allowed and forbidden transitions of interest for spectroscopic diagnostics in fusion plasmas

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Abstract
Transition probabilities for allowed and forbidden lines in neutral, singly ionized and doubly ionized tungsten are discussed in the present paper. For the electric dipole transitions, recommended values are proposed from a critical evaluation of the data available in the literature. For the magnetic dipole and electric quadrupole transitions, for which no data have been published so far, a new set of radiative rates has been obtained using a relativistic Hartree-Fock approach including core-polarization effects. The tables summarizing the compiled data are expected to be useful for plasma modelling in fusion reactors.

Recommended A-values for the most intense W I, W II and W III E1 lines

+ new decay rates for forbidden lines

The specific case of tungsten (W IV, W V, W VI)

Transition rates in W IV, W V and W VI

W IV : 105 experimentally known energy levels within the $5d^3$, $5d^26s$, $5d6s^2$, $5d^26p$ and $5d6s6p$ configurations (Kramida & Shirai, At. Data Nucl. Data Tables **95**, 305, 2009)

W V : 59 experimentally known energy levels within the $5d^2$, $5d6s$, $5d6p$, $6s6p$, $5d5f$ and $5d7p$ configurations (Kramida & Shirai, At. Data Nucl. Data Tables **95**, 305, 2009)

W VI : 14 experimentally known energy levels within the $5d$, $6s$, $6p$, $6d$, $7s$, $5f$, $5g$ and $6g$ configurations (Kramida & Shirai, At. Data Nucl. Data Tables **95**, 305, 2009)

Transition probabilities and oscillator strengths calculated for :

W IV lines in the range 143 – 990 nm (Enzonga Yoca *et al.*, J. Phys. B **45**, 035001, 2012)

W V lines in the range 83 – 1494 nm (Enzonga Yoca *et al.*, J. Phys. B **45**, 065001, 2012)

W VI lines in the range 38 – 1488 nm (Enzonga Yoca *et al.*, J. Phys. B **45**, 035002, 2012)

The specific case of tungsten (W IV, W V, W VI)

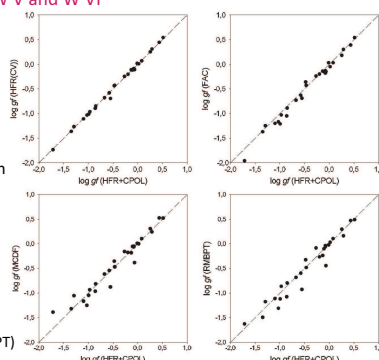
Transition rates in W IV, W V and W VI

No experimental radiative data to compare with !

Use of different and independent theoretical methods

HFR + CPOL compared with

- HFR including core-valence (HFR(CV))
- Flexible Atomic Code (FAC)
- Multiconfiguration Dirac-Fock (MCDF)
- Relativistic Many-Body Perturbation Theory (RMBPT)



Conclusions

New radiative data (lifetimes, transition probabilities, oscillator strengths) obtained for a large number of lines belonging to heavy neutral and lowly ionized atoms (fifth row, sixth row, lanthanides) [**~ 90 ions considered, ~ 700 lifetimes measured, ~ 100000 A-values calculated**]

Semi-empirical model based on the relativistic Hartree-Fock method including core-polarization effects

Good agreement between theoretical approach and experimental results

New data very useful in astrophysics, plasma physics, ...

Level structure still too poorly known for many ions to perform semi-empirical calculations (→ new term analyses needed)

Collaborations

Atomic structure calculations

Astrophysics & Spectroscopy, UMONS, Belgium (* also ULg, Liège, Belgium)

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Department of Physics, University of Brazzaville, Congo

(S. Enzonga Yoca)

Experimental measurements

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Department of Physics, Jilin University, Changchun, China

(Z. Dai, L. Han, Z. Jiang, P. Li, Z. Ma, G. Sun, S. You, J. Xu, W. Zhang, Y. Zhang)

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Faculty of Physics, Universidad Complutense de Madrid, Spain

(R. Mayo, M. Ortiz)

Thank you
for your
attention !